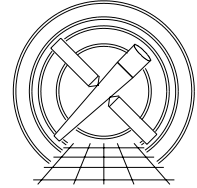




MIT  
Center for Space Research



Chandra X-Ray Center

## MEMORANDUM

January 15, 2005

**To:** Martin Elvis, SDS Group Leader  
**From:** Glenn Allen, SDS  
**Subject:** Computation of the Coordinates X, Y, and SKY\_1D and the Times of Arrival for Continuous-Clocking Event Data  
**Revision:** 4.5  
**URL:** <http://space.mit.edu/CXC/docs/docs.html#toa>  
**File:** `/nfs/cxc/h2/gea/sds/docs/memos/memo_acis_cc_calc_toa_4.5.tex`

The event times in ACIS Level 0 continuous-clocking (CC) event-data files are associated with the times events are read, not the times of arrival of the particles that deposit charge in the detector. Therefore, analyses of the event times to search for evidence of pulsations may be complicated by the time it takes to move charge from the nominal aim point to the location of the read-out electronics, the motion of the source on the detector as the telescope dither, the motion of the SIM relative to the telescope, and the use of the gratings. Analyses of the sky coordinates of continuous-clocking mode event data can be difficult if the effects of dither are not removed and if the source events are arbitrarily assumed to have CHIPY values of 512. To make it easier for users to analyze continuous-clocking mode data, `acis_process_events` (1) computes the times of arrival of events for a given source location, (2) removes the effects of dither and the SIM motion from the coordinates X and Y, (3) has the one-dimension image of the coordinates X and Y pass through the location associated with RA\_TARG and DEC\_TARG, and (4) computes the coordinate SKY\_1D (described below).

### 1 `acis_process_events`

The tool `acis_process_events` can create columns called `TIME_RO` and `SKY_1D` for continuous-clocking event data. The contents of the column `TIME_RO` correspond to the times events are read (i.e. are the same as the values of `TIME` in Level 0 event files). The contents of the column `TIME` are the estimated times of arrival of events assuming the events are from a source whose celestial coordinates are given by `RA_TARG` and `DEC_TARG`. Since it is not possible to discriminate between events associated with the source and background events, all events are handled in the same manner. The coordinate `SKY_1D` represents the distance of an event from the location of the source (in arcseconds) in the direction perpendicular to the readout direction.

### Additional Parameters

1. `calc_cc_times,b,h,"yes",,,` "Estimate the times of arrival for a CC-mode observation?"

## Input

1. An ACIS continuous-clocking event-data file that includes the columns TIME, CCD\_ID, and CHIPX (or CCDX).
2. The associated aspect solution file.
3. The right ascension (RA\_TARG) and declination (DEC\_TARG) of the observed source in J2000.0 coordinates.

## Output

1. An ACIS event-data file that includes the columns TIME\_RO (if specified as part of the parameter eventdef) and SKY\_1D.
2. If the value of the parameter calc\_cc\_times is “yes,” the header of the output file should contain the keyword HDUCLAS3 = ‘CC\_CORRECTED’ to confirm that the values in the column TIME are the estimated times of arrival instead of the read-out times. Otherwise, this keyword is excluded from the output file.

## Processing

1. Check for input errors: Verify that the input event files are continuous-clocking mode data files and contain columns named TIME (or TIME\_RO), CCD\_ID, and CCDX (or CHIPX). Verify that the keywords RA\_NOM, DEC\_NOM, RA\_TARG, DEC\_TARG and TIMEDEL exist. Verify that the input aspect solution file(s) exist.
2. Read the values of the keywords RA\_NOM, DEC\_NOM, RA\_TARG, DEC\_TARG, TIMEDEL and HDUCLAS3 (if it exists) from the header in the event-data file.
3. For event  $i$ , read the values of  $TIME_i$ ,  $TIME\_RO_i$  (if it exists),  $CCD\_ID_i$ ,  $CCDX_i$  (or  $CHIPX_i$ ), and  $CCDY_i$  (or  $CHIPY_i$ ). If the columns CHIPX and CHIPY do not exist, compute the values of  $CHIPX_i$  and  $CHIPY_i$ .
4. If the value of the parameter calc\_cc\_times is “yes,” the keyword HDUCLAS3 is “CC\_CORRECTED” and the column TIME\_RO does not exist in the input file, set  $TIME\_RO_i = TIME_i$ . See table 1 for a list of how to handle the other possible cases.
5. If the parameter calc\_cc\_times is “yes” or the parameter stop is “sky,” estimate the time of arrival  $t_i$  of event  $i$ . This estimate is only used to compute the value  $CHIPY_{TARG_i}$  (see step 7). It is not used to compute the time of arrival  $TIME_i$ . If event  $i$  is the first event in the input event-data file,

$$t_1 = TIME\_RO_1 - (512 + 1028) \times TIMEDEL, \quad (1)$$

where  $TIME\_RO_1$  is the read-out time of the first event. For all subsequent events,

$$t_i = TIME\_RO_i - (CHIPY_{TARG_{i-1}} + 1028) \times TIMEDEL, \quad (2)$$

where  $TIME\_RO_i$  is the read-out time of the  $i^{\text{th}}$  event and  $CHIPY_{TARG_{i-1}}$  is the CHIPY coordinate associated with RA\_TARG and DEC\_TARG at  $t_{i-1}$  (see step 7).

6. If the parameter calc\_cc\_times is “yes” or the parameter stop is “sky,” find the appropriate values of  $RA_i$ ,  $DEC_i$ , and  $ROLL_i$  associated with  $t_i$  in the aspect solution file(s).
7. If the parameter calc\_cc\_times is “yes” or the parameter stop is “sky,” use the values of  $RA_i$ ,  $DEC_i$ ,  $ROLL_i$ , RA\_TARG, DEC\_TARG, RA\_NOM and DEC\_NOM to compute the value of  $CHIPY_{TARG_i}$ , the CHIPY coordinate associated with RA\_TARG and DEC\_TARG at time  $t_i$ . The value of  $CHIPY_{TARG_i}$  is used to compute the values of  $TIME_{i,X_i}$ ,  $Y_i$  and  $SKY\_1D_i$ .

Table 1. Possible cases

Parameter calc_cc_times	infile			outfile		
	Keyword HDUCLAS3	Column TIME	Column TIME_RO	Keyword HDUCLAS3	Column TIME	Column TIME_RO*
yes <sup>†</sup>	Not CC_CORRECTED	Exists	Doesn't exist	CC_CORRECTED	Input TIMES + Adjustments	Input TIMES
yes <sup>†</sup>	Not CC_CORRECTED	Exists	Exists	CC_CORRECTED	Input TIME_ROs + Adjustments	Input TIME_ROs
yes <sup>†</sup>	CC_CORRECTED	Exists	Doesn't exist	CC_CORRECTED	Input TIMES	Zeroes
yes <sup>†</sup>	CC_CORRECTED	Exists	Exists	CC_CORRECTED	Input TIME_ROs + Adjustments	Input TIME_ROs
no	Not CC_CORRECTED	Exists	Doesn't exist	Doesn't exist	Input TIMES	Zeroes
no	Not CC_CORRECTED	Exists	Exists	Doesn't exist	Input TIMES	Input TIME_ROs
no	CC_CORRECTED	Exists	Doesn't exist	CC_CORRECTED	Input TIMES	Zeroes
no	CC_CORRECTED	Exists	Exists	CC_CORRECTED	Input TIMES	Input TIME_ROs

<sup>†</sup> If the parameter acaoffile=NONE or the file(s) specified by acaoffile do not exist, then produce a warning message and reset the parameter calc\_cc\_times from “yes” to “no.”

\* If the column is specified as part of the output using the parameter “eventdef.”

8. If the parameter calc\_cc\_times is “yes,” compute the time of arrival of event  $i$ :

$$\text{TIME}_i = \text{TIME\_RO}_i - (\text{CHIPY}_{\text{TARG}_i} + 1028) \times \text{TIMEDEL}.$$

Note that the contents of the column TIME contain the estimated times of arrival instead of the read-out times. The column TIME\_RO contains the read-out times.

9. If the parameter stop is “sky,” compute the values of  $X_i$  and  $Y_i$  using the values of  $\text{TIME}_i$ ,  $\text{CCD\_ID}_i$ ,  $\text{CHIPX}_i$  and  $\text{CHIPY}_{\text{TARG}_i}$  (not  $\text{CHIPY}_i$ ). The resulting sky coordinates are free of the effects of dither and the motion of the SIM. The sky image is a thin line (only a few pixels wide) that passes through RA\_TARG and DEC\_TARG.
10. If the parameter stop is “sky,” compute the value of SKY\_1D.
11. Repeat steps 3 through 10 for each event in the input event-data file.
12. Write the results to the output event-data file. Include the column TIME\_RO and the keyword HDUCLAS3, if appropriate (see table 1).

## 2 Acknowledgements

The development of the algorithms to compute the sky positions and times of arrival of the events relied heavily on the assistance of Ian Evans, Peter Ford, Kenny Glotfelty, David Huenemoerder, Jonathan McDowell, Herman Marshall, Joe Masters, Arnold Rots, Divas Sanwal, and Allyn Tennant. Jonathan McDowell produced much of the code to compute the sky positions of the events and Herman Marshall and Allyn Tennant provided a great deal of help with the algorithm to compute the times of arrival.